
Impact of Imaging Optics Quality on Actinic Mask Blank Inspection

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Current Organization:

1: ASET EUV Process Technology Laboratory

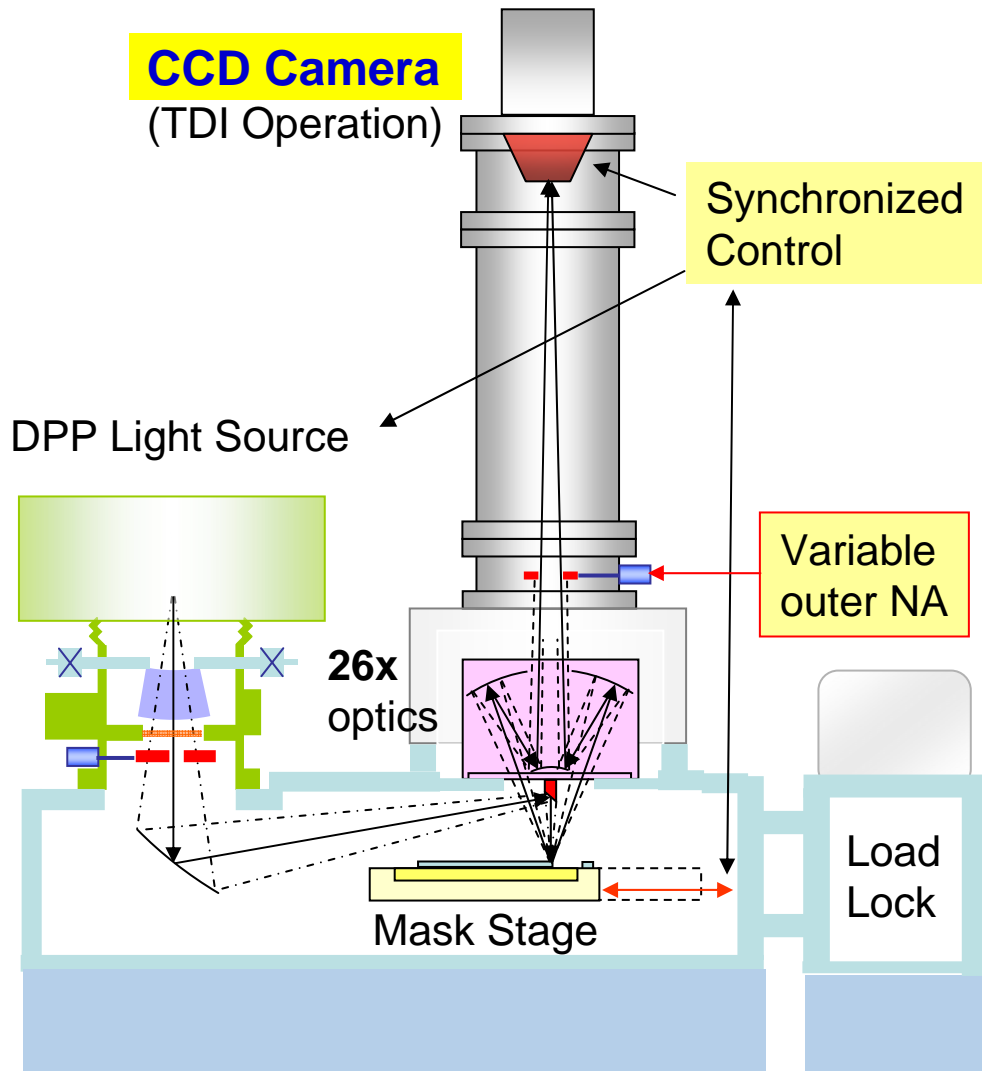
2: MIRAI-Selete, 3: ASRC-AIST



Outline

- Introduction
 - Prototype design
 - Sensitivity-limiting factors
- Impact of imaging optics quality
 - Defect image with 26x optics
 - PSF extraction from defect signal
 - Dark-field flare evaluation
- Discussion
 - Comparison with optics MSFR
 - Requirement for optics quality
- Summary

Full-field Prototype Design

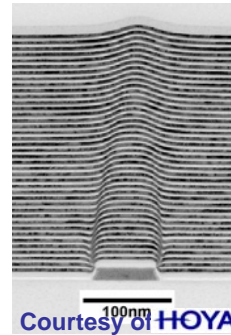
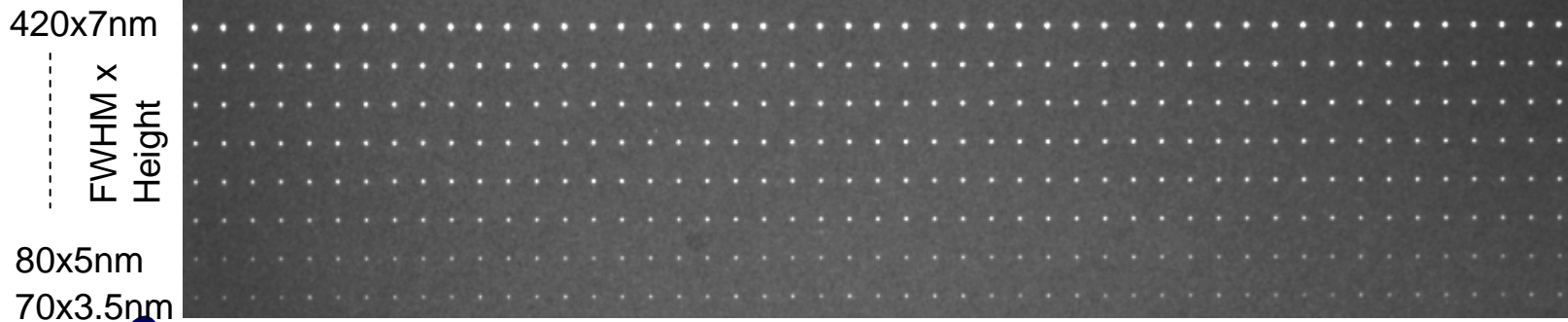
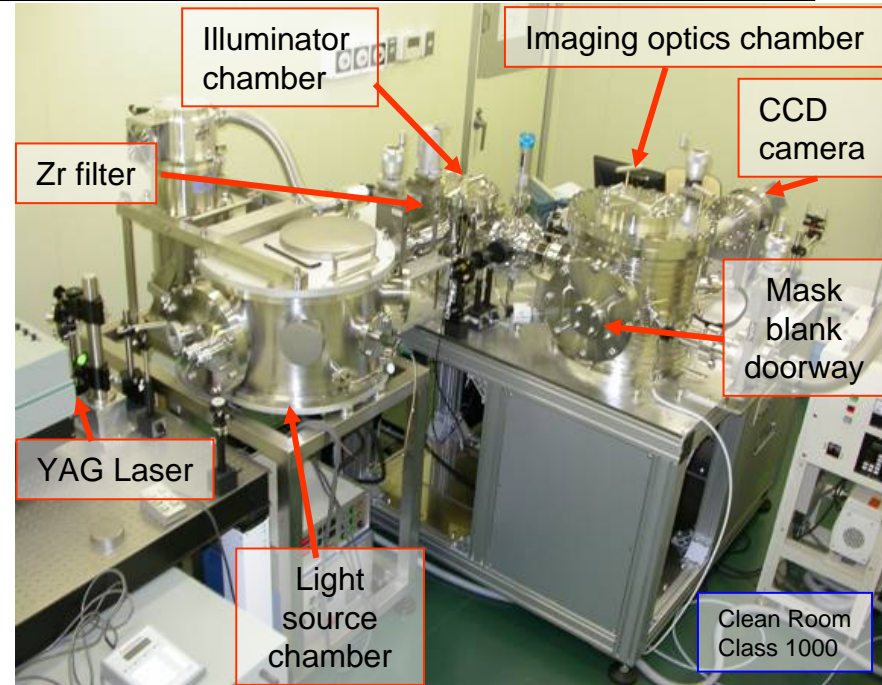
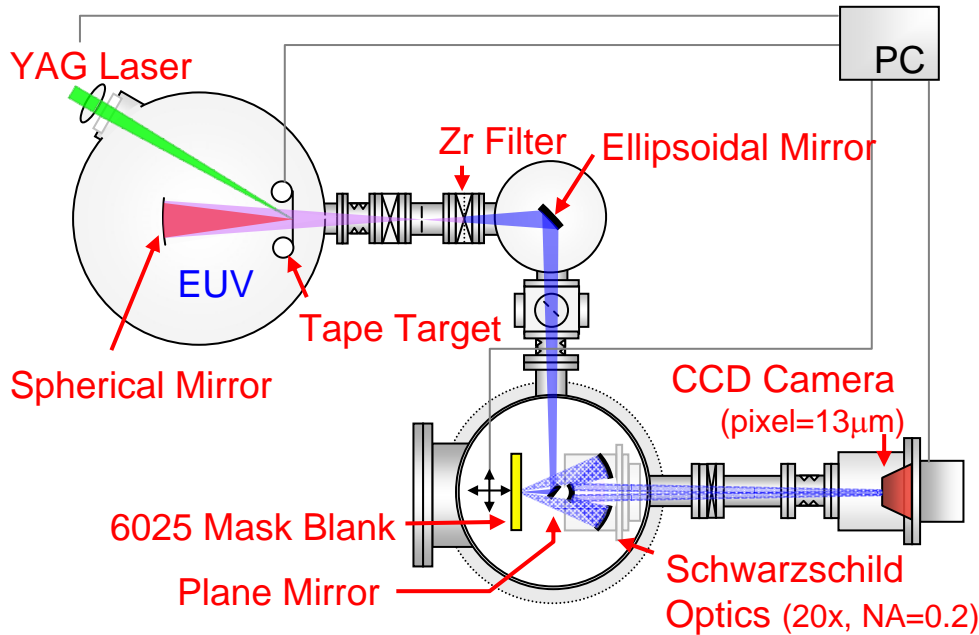


Model	MIRACL-1000 ⁽¹⁾
Specification	
Objects	6025 EUVL Mask Blanks
Sensitivity	$< 40\text{nm}^w \times 2\text{nm}^h$ ⁽²⁾
Throughput	2hrs/blank
Pos. Accuracy	$< 250\text{nm}$
Configuration	
Light Source	wavelength=13.5nm, DPP, Every pulse Triggerble
Illumination	Critical Illumination, Ellipsoidal mirror + Plane mirror, Illum. area $> 0.5\text{mm}^2$
Imaging Optics	26x Schwarzschild Optics, Inner NA = 0.1, Outer NA = 0.2~0.3
Sensor	Back-illuminated CCD, TDI Operation, Synchronized with light source pulses
Stage	Continuous move, Interferometer feedback
Software	Automatic defect classification, Size inference, Position identification
Add-on Func.	Angular dist measurement with variable aperture for shape inference
Alignment	Focus / Position alignment function by fiducal mark
Interface	SMIF-capable, EUV mask handling standard compliant

(1: Maskblank Inspection for Reflective multilayer by Actinic Light)

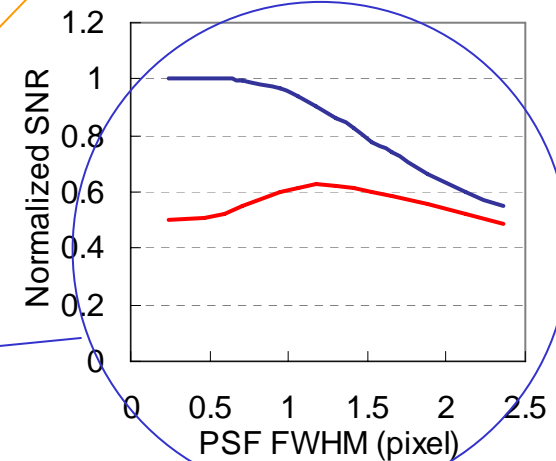
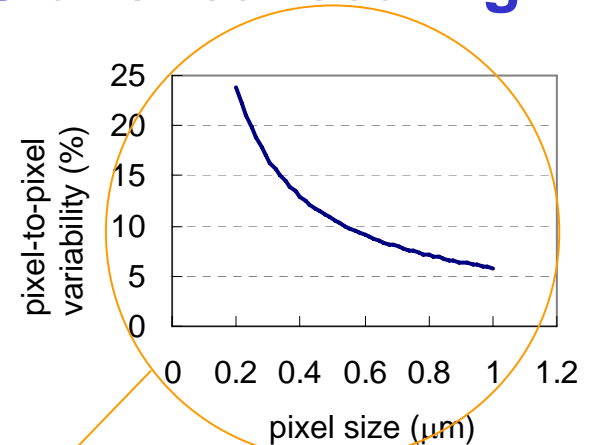
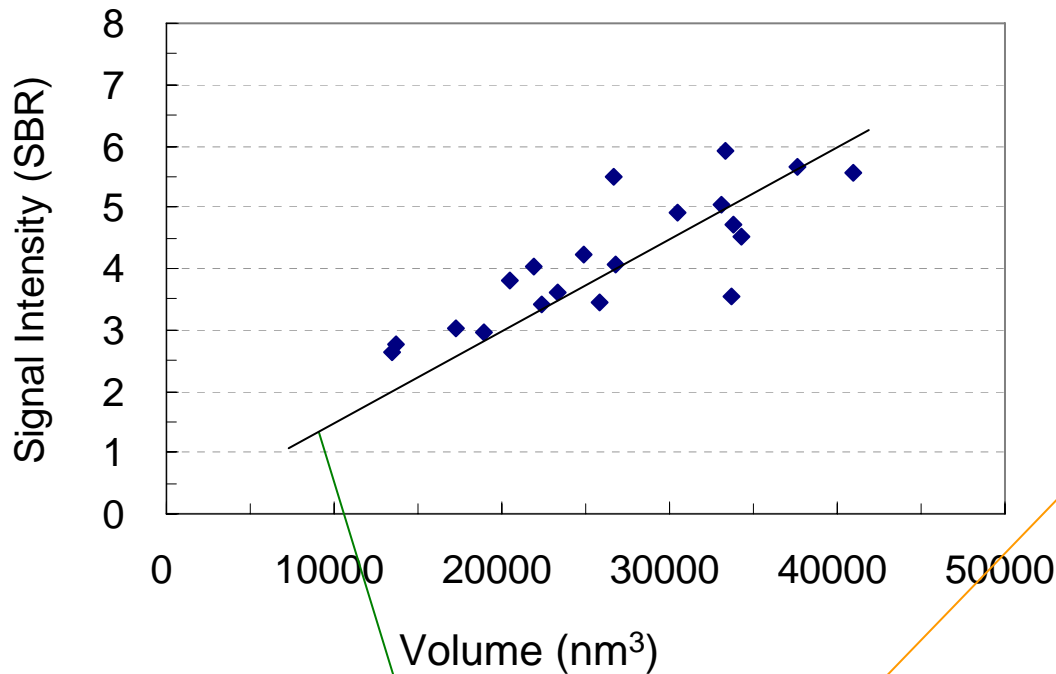
(2: in use of ultra smooth blank)

POC Tool & Inspection Image



Sensitivity Derivation

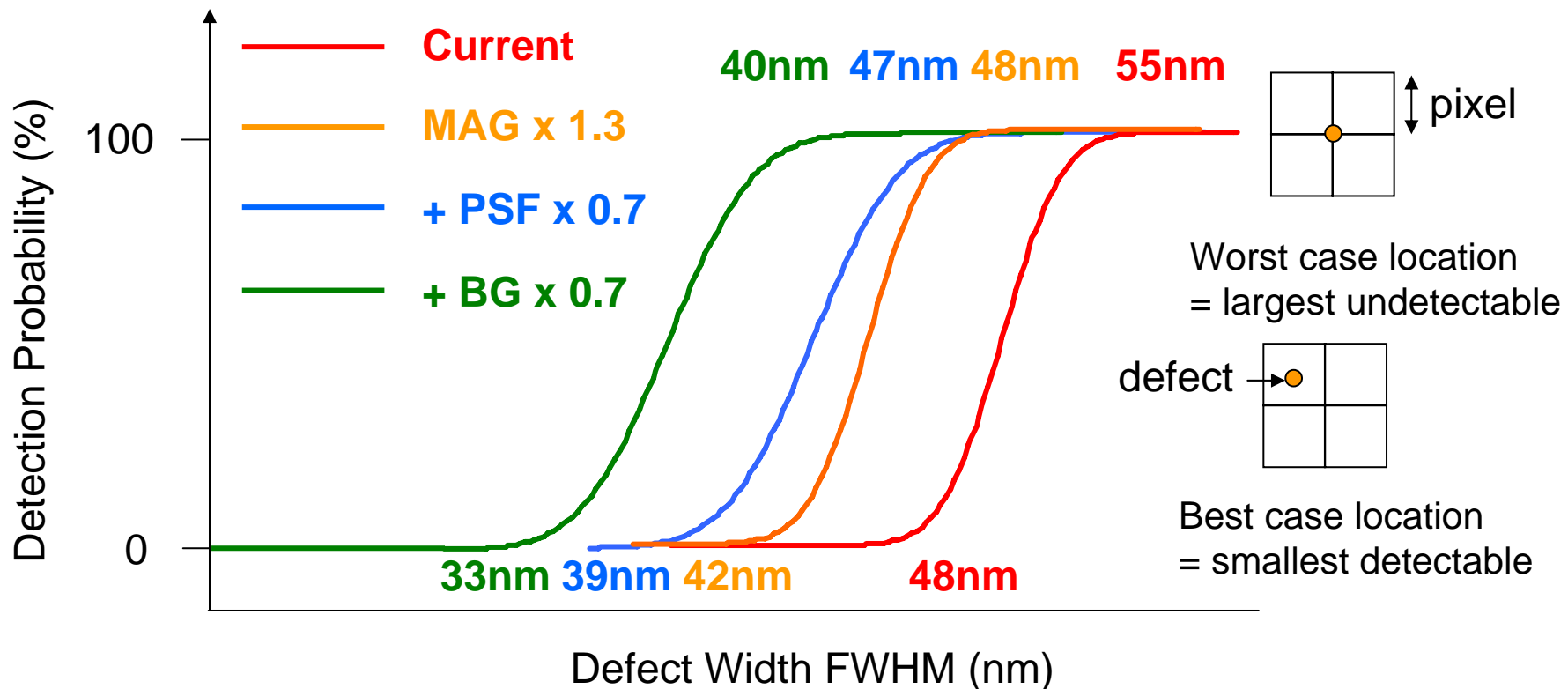
Extrapolation from Experiments + Statistical scaling



$$SNR = \frac{S(\text{volume})}{B(\text{roughness, pixel}) \cdot \sigma_r(\text{pixel})} \cdot \frac{\xi(\text{PSF})}{f(\text{PSF})}$$

Sensitivity Prediction

FWHM of 2nm-high Detectable Defects at SNR>7

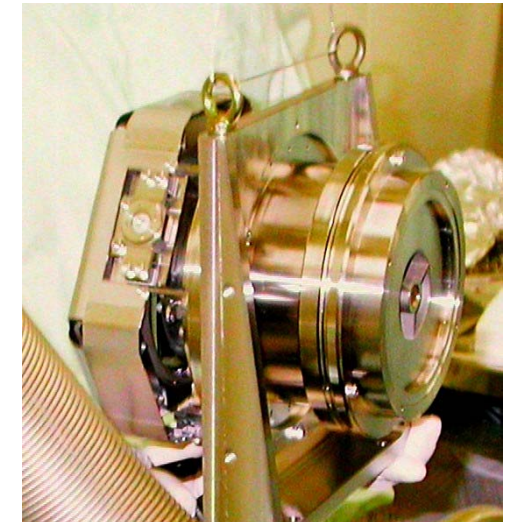
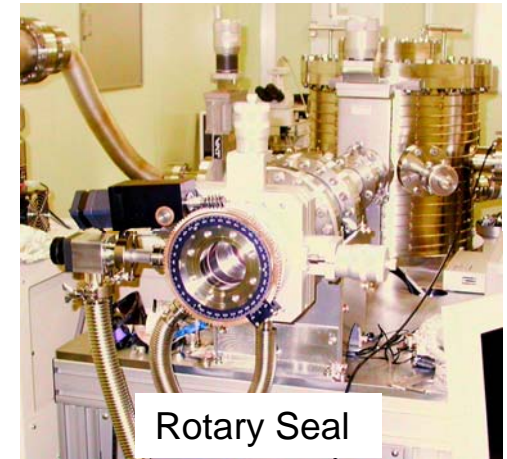
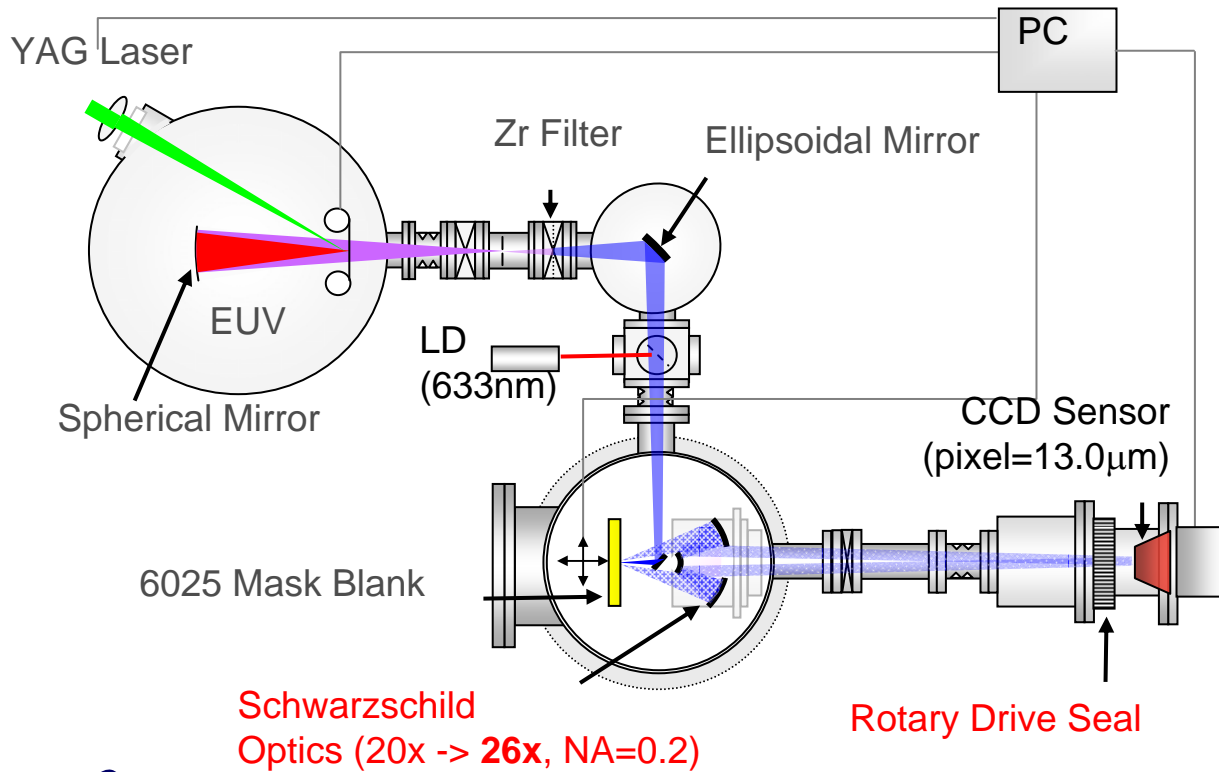


Blank Roughness Reduction is Essential !

POC Tool Upgrade

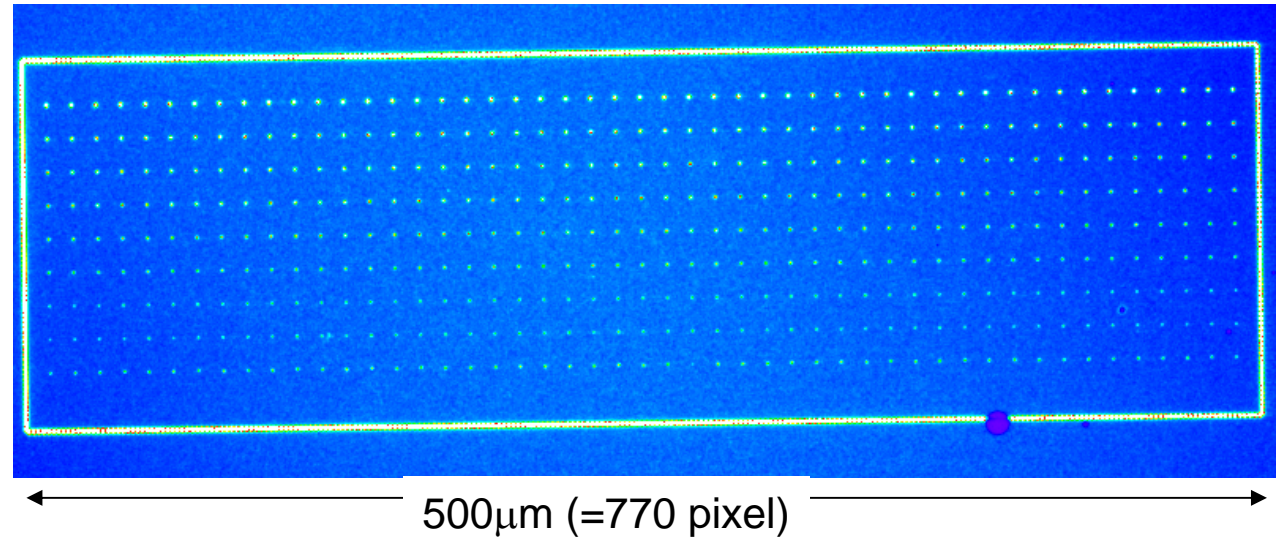
Latest Upgrade

- Schwarzschild Optics: 20x -> 26x
 - best effort-based quality within limited time frame
- CCD Rotary Stage

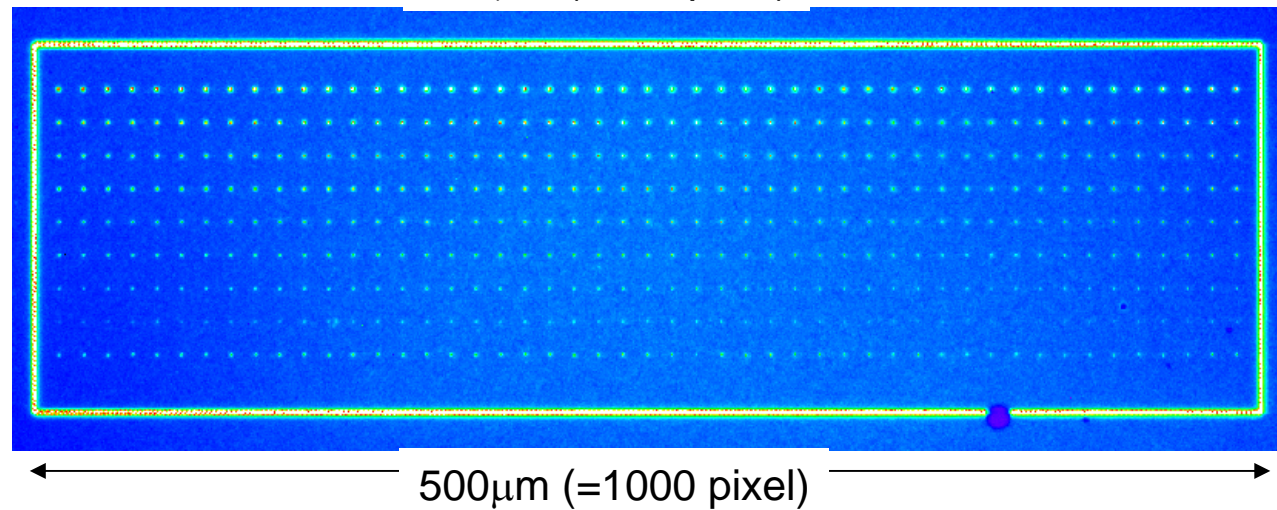


Inspection Image Comparison

20x, May'05



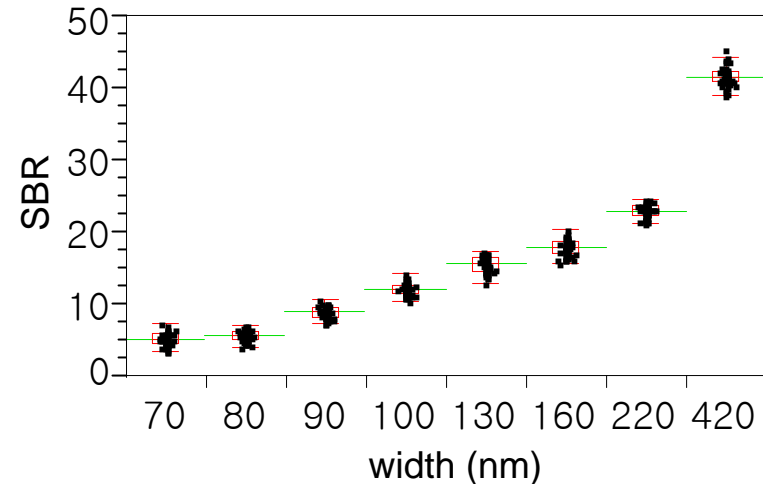
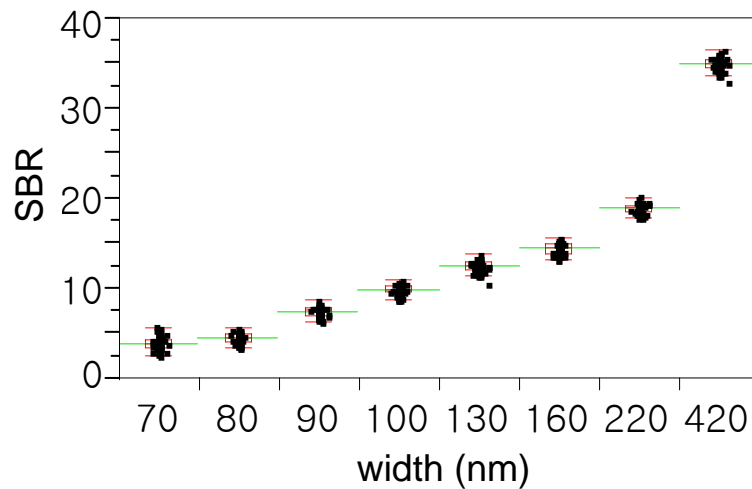
26x, Jan'06



Gross Intensity Comparison

20x (5x5 pixel)

26x (6.5x6.5 pixel)



$$SBR = \sum_{i,j}^{5 \times 5 \text{ or } 6.5 \times 6.5} \frac{A_{i,j} - \langle BG \rangle_{\text{Local}}}{\langle BG \rangle_{\text{Local}}}$$

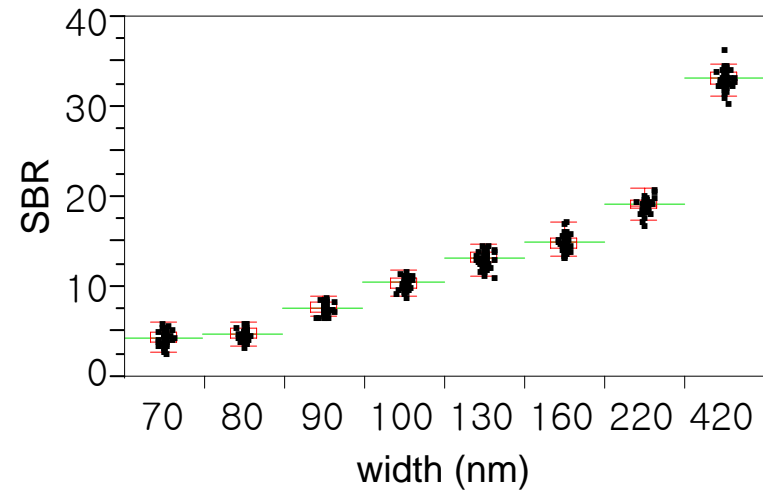
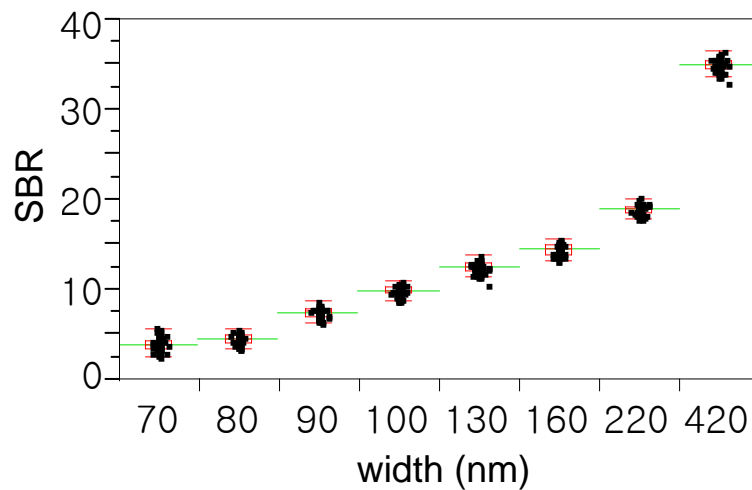
Theoretical Prediction: 1.69x improvement if optic finish is the same
 Experiment: 1.25x improvement

EUV after MSFR loss of 26x optics can be only 74%(=1.25/1.69) of 20x optics.

Normalized Intensity Comparison

20x (5x5 pixel)

26x (5x5 pixel)



$$SBR = \sum_{i,j}^{5 \times 5} \frac{A_{i,j} - \langle BG \rangle_{Local}}{\langle BG \rangle_{Local}}$$

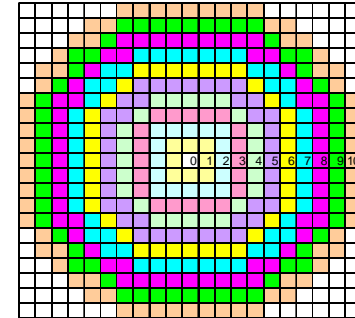
Fair comparison metric in working sensitivity

Prediction: 1.69x improvement **provided that PSD improves by 1.3x**

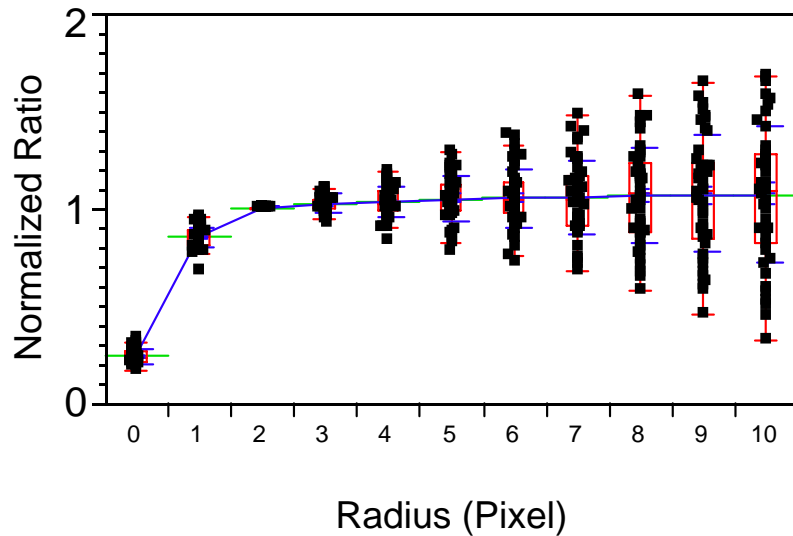
Experiment: 1.06x improvement -> virtually no improvement

Encircled Energy Comparison

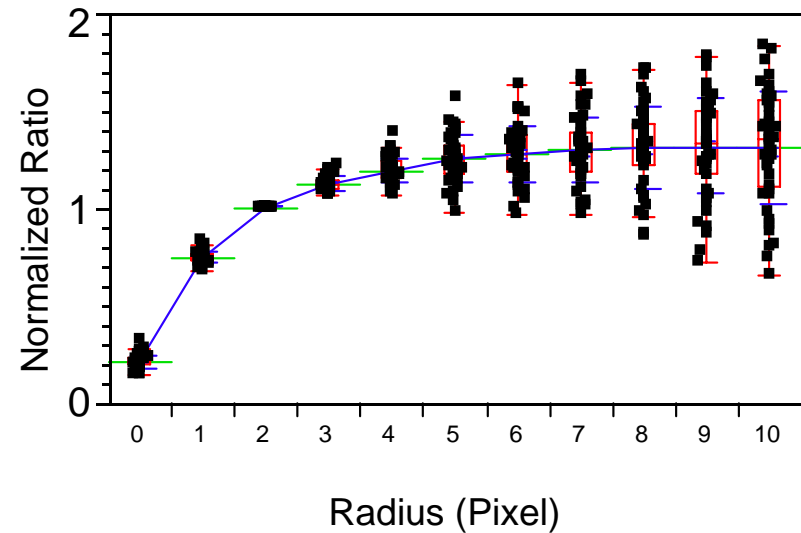
- Integrated PSF extraction from images



20x optics



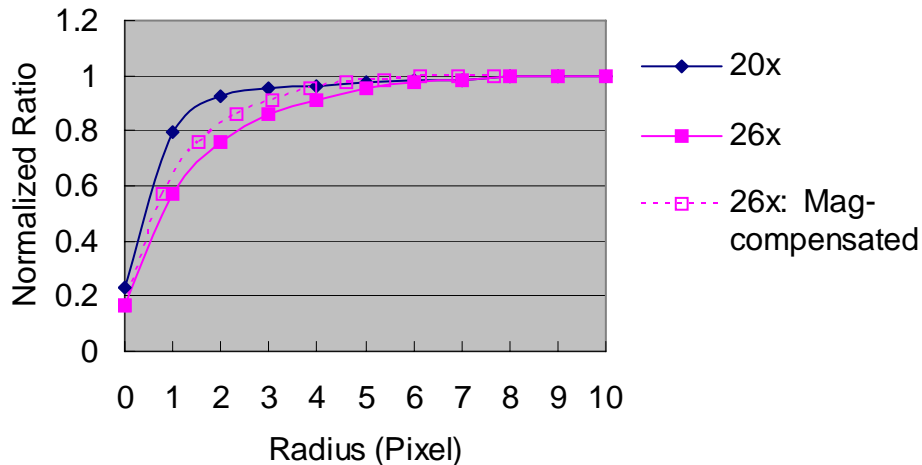
26x optics



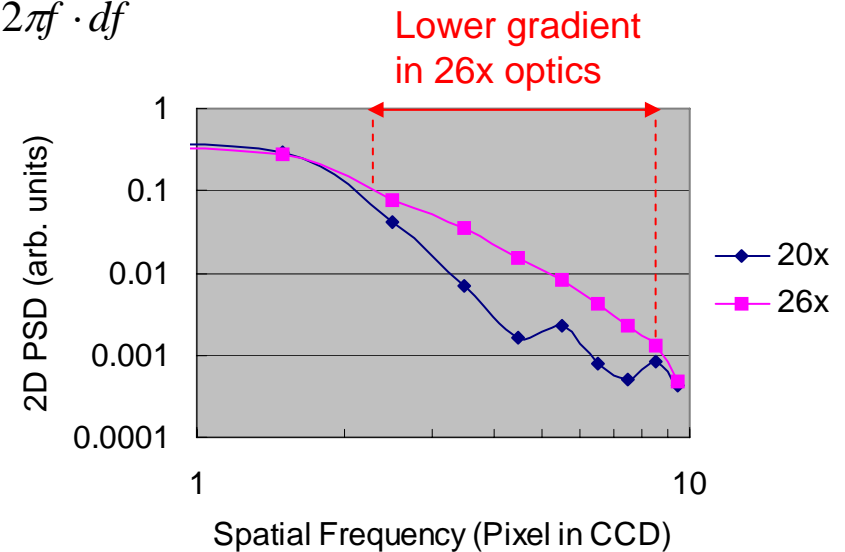
Translation to PSD

Comparison of Average Integrated PSF

$$I(P) = \int_0^P PSD(f) 2\pi f \cdot df$$



Translated PSD Curves



Quality of 26x optics worse than 20x.

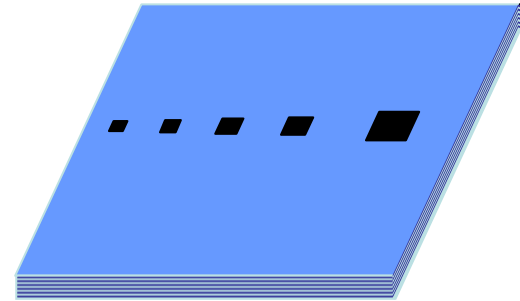
Lower gradient in 2~8 pixels



Roughness in spatial frequency of :
 700um~2.9mm in M1(concave) and/or
 150um~600um in M2 (convex) can be
 worse in 26x optics

Evaluation of shorter MSFR

- Test mask:
 - Clear filed mask with various sizes of absorber pads
- Evaluation procedure
 - Get darkfield image and measure “flare” from surrounding background intensity at pad center



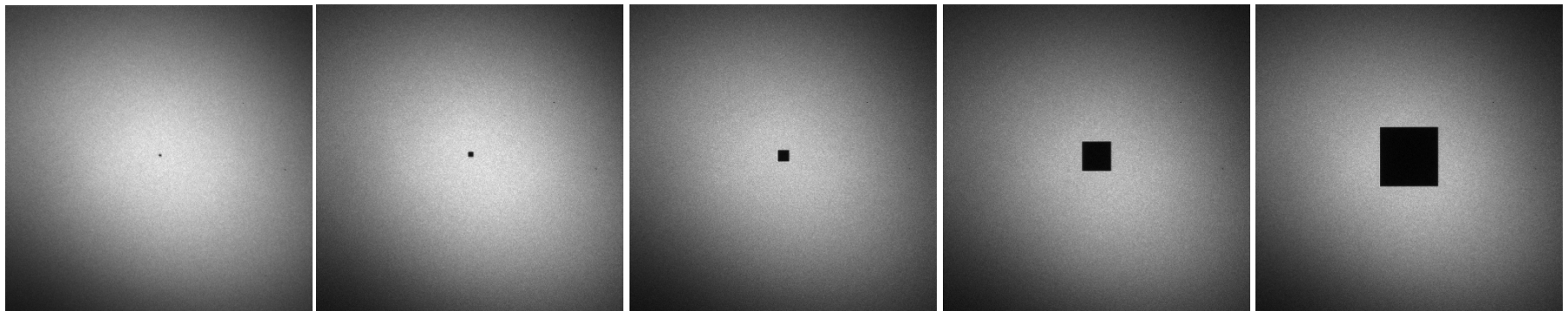
Pad: 5um

10um

20um

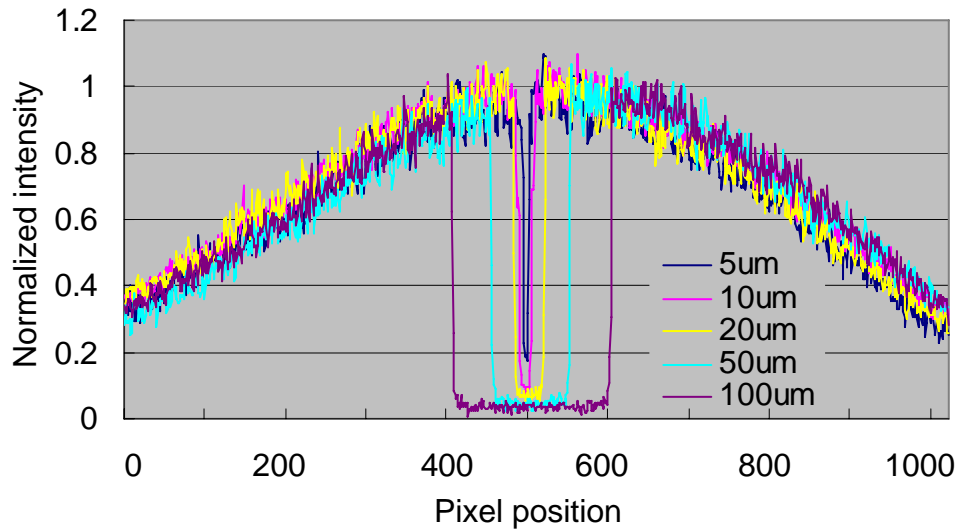
50um

100um



← 500um →

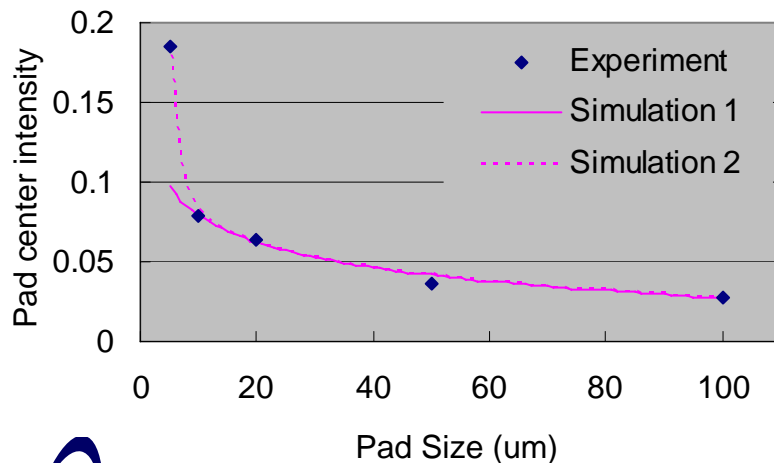
Comparison with model PSD



$$F_{center}(R_{PAD}) = \int_{R_{PAD}}^{\infty} I(r) \cdot PSF(r) \cdot 2\pi r dr$$

$$I(r) = \exp\left(-\frac{r^2}{R_{ill}^2}\right) : \text{Illumination distribution}$$

$$PSF(r) = \frac{16\pi^2}{\lambda^4 \cdot \rho_2^2} \cdot \left(PSD_2\left(\frac{r}{\lambda \cdot \rho_2}\right) + \alpha^2 \cdot PSD_1\left(\frac{\alpha \cdot r}{\lambda \cdot \rho_2}\right) \right)$$



Simulation 1:

both PSD $\sim 1/f^{2.1}$,
 (M1: 60um~1.2mm, M2: 12~240um),
 MSFR=0.29nm

Simulation 2:

PSD1(M1) $\sim 1/f^5$ for 400um~1.2mm
 Others are same with Sim 1

Consistent with PSF characterization

Discussion

- Supplier data turn out to corroborate worse MSFR in 26x optics

MSFR (nm RMS: 1um~1mm)

	20x	26x
M1(convex)	0.36	0.41
M2(concave)	0.26	0.57

- Image data analysis suggests the worse area is in longer spatial period range in MSFR range and beyond into figure error range.
- Preferably compared with each PSD if available

- Less-than-anticipated improvement in sensitivity well explained by worse optics quality.
 - Not a counterproof of sensitivity improvement path by magnification.
 - Rather, more-than-predicted improvement possible as 20x optics was still below state-of-the-art quality.
- Target spec for 26x optics are:
 - Assuming >90% within radius of 1 pixel (13um)

$$\left(\frac{4\pi}{\lambda}\right)^2 \cdot \int_{f_{\min}}^{f_{\max}} PSD(f) \cdot 2\pi f \cdot df = \left(\frac{4\pi\sigma}{\lambda}\right)^2 < 0.05$$

$$\therefore \sigma < 0.24\text{nm}$$

for roughness period integration range of:

7um ~ **7mm** in M1 and 1.4um ~ **1.4mm** in M2

Summary

- Upgrade of projection optics from 20x to 26x implemented.
- Less-than-anticipated improvement in sensitivity attributed to worse MSFR to short range figure errors of mirrors by image data analysis.
- More improvement in sensitivity than previous prediction expected by using state-of-the-art quality mirrors.

Acknowledgments

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 - T. Shoki, K. Yamashiro, Y. Usui and O. Nagarekawa of HOYA Corporation for their fabrication of the programmed defect mask blank.
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The logo for HOYA Corporation, featuring the word "HOYA" in a bold, blue, sans-serif font.